# On the minimization properties of the **Tisserand systems**

Alberto Escapa<sup>1</sup>, Tomás Baenas<sup>1</sup>, José Manuel Ferrándiz<sup>1</sup>, & Juan Getino<sup>2</sup>

<sup>1</sup>Dept. Applied Mathematics, University of Alicante, P. O. 99, Alicante 03080, Spain <sup>2</sup>Dept. Applied Mathematics, University of Valladolid, Valladolid 47011, Spain

(5)

Alberto.Escapa@ua.es

#### **1. Introduction**

\* HER FACITE ELS

 $\Lambda$  [HEN a collection of material particles S experience relative displacements, it is no possible to define unambiguously a rotational motion of the set. In these situations it is assigned to  $\mathcal{S}$  a reference system Oxyz (the "body axes") with origin in its barycenter O and connected with it in a prescribed way.

From Eqs. (1) and (3), the deformation kinetic energy can be written as (Escapa 2011)

$$\mathcal{T}_{ ext{def}}\left(ec{\omega}
ight) = \mathcal{T} - ec{L}ec{\omega} + rac{1}{2}ec{\omega}\,ec{\omega}\,ec{\omega},$$

B<sup>Y</sup> doing so, the rotation of the particles is identified with the ro-tation of the body axes with respect to some inertial, or quasiinertial, reference system OXYZ. This rotation admits a precise definition in mathematical terms.

THERE are different possibilities to connect the body axes Oxyzwith the considered set of particles (Munk & McDonald 1960). From the point of view of simplifying the equations of motion, one convenient method is to employ the so-called Tisserand systems (Tisserand 1891).

# 2. Tisserand systems

O introduce Tisserand systems, let us write the velocity, relative to OXYZ, of a particle of S with position  $\vec{x_i}$  and mass  $m_i$  as

> $\vec{V}_i = \vec{\omega} \times \vec{x}_i + \vec{v}_i(\vec{\omega}).$ (1)

where  $\mathcal{T}$  is the kinetic energy of  $\mathcal{S}$ . Hence, for an arbitrary vector  $\vec{\lambda}$ different from  $\vec{0}$ , we have

$$\mathcal{T}_{def}\left(\vec{\omega}+\vec{\lambda}\right) = \mathcal{T}_{def}\left(\vec{\omega}\right) - \vec{L}\,\vec{\lambda} + \vec{\lambda}\,\mathbf{I}\,\vec{\omega} + \frac{1}{2}\vec{\lambda}\,\mathbf{I}\,\vec{\lambda}.$$
 (6)

If we consider condition (a), defining the angular momentum of the system  $\vec{L}$ , in Eq. (6), we get

$$\mathcal{T}_{def}\left(\vec{\omega}_T + \vec{\lambda}\right) - \mathcal{T}_{def}\left(\vec{\omega}_T\right) = \frac{1}{2}\vec{\lambda} \,\mathbf{I}\,\vec{\lambda}.$$
(7)

Since the matrix of inertia is definite positive, we have that

 $\frac{1}{2}\vec{\lambda} \, \mathbf{I} \, \vec{\lambda} > 0, \ \vec{\lambda} \in \mathbb{R}^3, \ \vec{\lambda} \neq \vec{0}.$ (8)

Therefore, Eq. (7) implies that  $\mathcal{T}_{def}(\vec{\omega})$  takes its minimum at  $\vec{\omega}_T$ , i.e., condition (b).

## 4. Tisserand systems evolution

THE angular velocity  $\vec{\omega}_T$ , considered as a known function of time, determines the rotational kinematics of the body axes, but not

The vector  $\vec{\omega}$  is, at this stage, arbitrary and common for the set S. In contrast  $\vec{v}_i(\vec{\omega})$ , the deformation velocity (Moritz & Mueller 1987), depends on the material particle *i* and the particular choice of  $\vec{\omega}$ .

Tisserand systems can be defined by any of the following conditions that fix  $\vec{\omega}$  to a particular value  $\vec{\omega_T}$ :

(a) The angular momentum of S

 $\vec{L} = \sum_{i \in \mathcal{S}} m_i \left( \vec{x}_i \times \vec{V}_i \right)$ (2)

(3)

(4)

is  $\vec{L} = |\vec{\omega_T}|$  (Tisserand 1891), where I is the matrix of inertia of S (b) The kinetic energy of S associated to the deformation velocity

$$\mathcal{T}_{\mathsf{def}}(\vec{\omega_T}) = \frac{1}{2} \sum_{i \in \mathcal{S}} m_i \vec{v_i} (\vec{\omega_T})^2$$

is minimum (Jeffreys 1976)

(c) The angular momentum of  $\mathcal{S}$  related with the deformation velocity

 $\vec{h}(\vec{\omega_T}) = \sum m_i \left[ \vec{x}_i \times \vec{v}_i(\vec{\omega_T}) \right]$ 

its orientation in an univocal manner (Tisserand 1891).

Specifically, from the components of  $\vec{\omega}_T$  in the OXYZ system, we can construct the skew-symmetric matrix

$$\Sigma_T(t) = \begin{pmatrix} 0 & -\omega_{TZ}(t) & \omega_{TY}(t) \\ \omega_{TZ}(t) & 0 & -\omega_{TX}(t) \\ -\omega_{TY}(t) & \omega_{TX}(t) & 0 \end{pmatrix}.$$
 (9)

It allows defining a rotation matrix R(t) that brings the OXYZ system to the body axes through (Wintner 1941)

$$\Sigma_T(t) = \frac{d\mathsf{R}^T}{dt}\mathsf{R},\tag{10}$$

where the superscript T denotes the transpose of a matrix. The solution of this linear differential equation is given by

$$\mathsf{R}(t) = \mathsf{R}(t_0) \exp\left(-\int_{t_0}^t \mathsf{\Sigma}_T(s) ds\right), \tag{1}$$

 $R(t_0)$  providing the numerical value of R(t) at the epoch  $t_0$ . In this way, besides any of the conditions (a), (b), or (c), the specifi-

 $i \in \mathcal{S}$ 

is the null vector (Tisserand 1891)

## 3. Equivalence of the conditions

**T**HE former characterizations turn out to be equivalent, that is to say, (a)  $\Rightarrow$  (b), (b)  $\Rightarrow$  (c), and (c)  $\Rightarrow$  (a). The second and third implications are detailed in the existing literature (e.g., Moritz & Mueller 1987). Let us focus on the first one.

Session 4. Earth's rotation and geodynamics

cation of a particular Tisserand system requires providing explicitly the initial orientation of the body axes relative to OXYZ.

#### References

• Escapa, A.: Celest. Mech. Dyn. Astr., Vol. 110, 99–142, 2011 • Jeffreys, H.: The Earth. Cambridge University Press, 1976 • Moritz, H. & Mueller, I.: Earth Rotation. Frederic Ungar, 1987 • Tisserand, F. F.: Traité de Mécanique Céleste, Vol. 2. Gauthier-Villars, 1891

 Wintner, A.: The Analytical Foundations of Celestial Mechanics. Princeton University Press, 1941

Journées "Systèmes de référence spatio-temporels", St. Petersburg, Russia, 22-24 September 2014