

Triaxial Earth's rotation: Chandler wobble, free core nutation and diurnal polar motion

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1 Introduction

For a biaxial elastic Earth, the free motions CW and FCN are circular motions.

When non-rotational-symmetric effects are taken into consideration, these free motion would become elliptical ones rather than circular ones.

An elliptical motion can be decomposed into two circular motions with opposite frequencies. So in the frequency domain, for every free motion, CW or FCN, there are two resonant frequencies. If the excitation function is very near to these two resonant frequencies, resonance will happen.

In this study, we used a matrix based method to study the influence of triaxiality on the frequencies of CW and FCN and on the diurnal polar motion.

2 Influence of triaxiality on frequencies of CW and FCN

To estimate the frequencies of CW and FCN, we treat the Euler-Liouville's equation as a set of linear differential equations with constant coefficients. Then the frequencies can be derived by using the theory of solution of a set of linear differential equations with constant coefficients.

We find that the influence of triaxiality on CW is very small, which is in accordance with the previous studies.

When it comes to the influence of triaxiality on FCN, we find that our results suggest that considering triaxiality of the core would make the FCN period shorter, which is in contradiction with the conclusion given by van Hoolst and Dehant (2002).

However, the theoretical system used by van Hoolst and Dehant(2002) is different from the theoretical system used in this study, which is the same as that used by Mathews et al., (2002) and that used by Chen and Shen (2010). The lateral inhomogeneity of the Earth is modelled as triaxial buldge in the study of van Hoolst and Dehant(2002), which means that the expression of FCN contains contribution of lateral inhomogeneity.

After we consider the contribution of lateral inhomogeneity, the period of FCN is also increased as shown in table 1.

Table 1 Influence of triaxiality on the period of FCN

	Biaxial Core	Triaxial Core
van Hoolst and Dehant (2002) ^b	429.613213	0.018699
This study (triaxiality alone)	430.941420	-0.007421
This study (triaxiality+lateral inhomogeneity)	430.941420	0.004864

3 Resonance at negative FCN

We solve the polar motion equation for a two-layer triaxial Earth model. The wobble can be decomposed into two parts: the same frequency ψ^o response m^s that is excited by excitation process ψ^s of same frequency and opposite frequency response m^o that is excited by excitation process of opposite frequency. The resonance can be shown in Figure 1.

To find out how large the resonances shown in figure 1 would be, we use the ocean tide model C of Chao(1996) as the excitation function. We focus on the resonance strength at negative FCN the resonance at negative FCN(1d) is only slightly smaller than resonance at FCN (1a). The results are shown in Table 2.

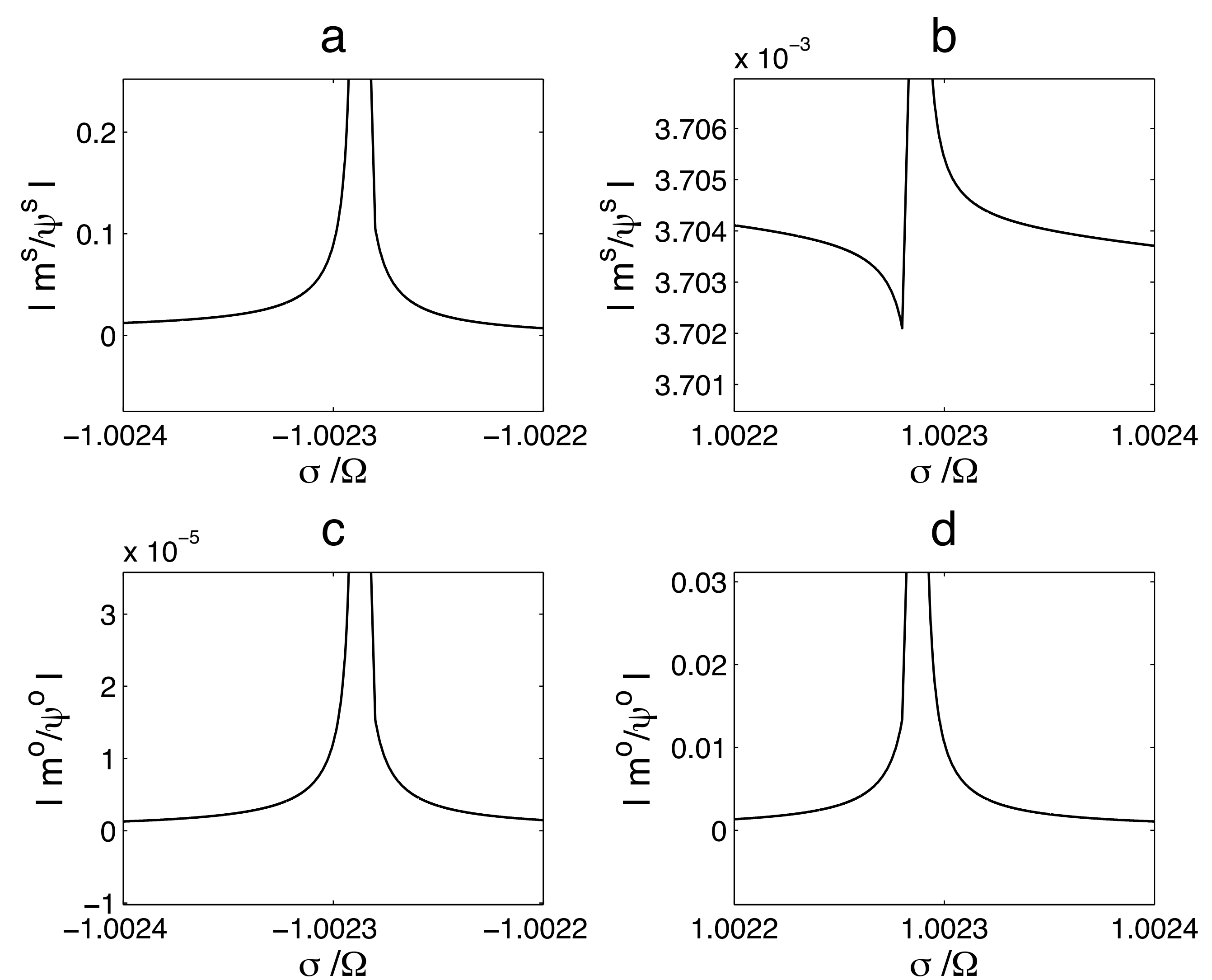


Figure 1 Resonances near FCN (a,c) and negative FCN(b,d). The resonances in (b,c,d) won't exist if the Earth mode is biaxial.

Table 2 Difference of polar motion amplitude between biaxial case and triaxial case. The closer the excitation is to the negative FCN, the larger the difference between the biaxial case and triaxial case.

Tide	Frequency (cpsd)	Triaxial-biaxial(μ s)
Q1	0.891	-0.002
O1	0.927	-0.018
P1	0.995	-0.089
K1	1.000	-0.954
Negative FCN	1.002288790	

4 Conclusion

Triaxiality would alter the frequencies of CW and FCN by a small amount. Triaxiality also makes the free motion of CW and FCN become elliptical motions rather than circular motions as suggested by biaxial Earth model.

Since an elliptical motion can be decomposed into two circular motion, there is an extra resonance frequency at complex Fourier's spectrum.

At negative FCN frequency, we find that this resonance would bring a difference of the order of 1μ s.

References

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